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# Non-convex anisotropic surface energy and zero mean curvature surfaces in the Lorentz-Minkowski space 

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#### Abstract

An anisotropic surface energy functional is the integral of an energy density function over a surface. The energy density depends on the surface normal at each point. The usual area functional is a special case of such a functional. We study stationary surfaces of anisotropic surface energies in the euclidean three-space which are called anisotropic minimal surfaces. For any axisymmetric anisotropic surface energy, we show that, a surface is both a minimal surface and an anisotropic minimal surface if and only if it is a right helicoid. We also construct new examples of anisotropic minimal surfaces, which include zero mean curvature surfaces in the three-dimensional LorentzMinkowski space as special cases.


Keywords. anisotropic, mean curvature, minimal surface, zero mean curvature surface, LorentzMinkowski space, Wulff shape

## 1. Introduction

Let $\gamma: \Omega \rightarrow \mathbf{R}_{+}$be a positive $C^{\infty}$ function on a nonempty open set $\Omega$ of the two-dimensional unit sphere $S^{2}:=\{X \in$ $\left.\mathbf{R}^{3} ;|X|=1\right\}$. Let $X: \Sigma \rightarrow \mathbf{R}^{3}$ be an immersion from a two-dimensional oriented connected compact $C^{\infty}$ manifold $\Sigma$ (with or without boundary) to the three-dimensional euclidean space $\mathbf{R}^{3}$. Denote by $\nu=\left(\nu_{1}, \nu_{2}, \nu_{3}\right): \Sigma \rightarrow S^{2}$ the unit normal along $X$ (in other words, the Gauss map of $X)$. If $\nu(\Sigma) \subset \Omega$, we say that $X$ is compatible with $\gamma$ and we define the following functional.

$$
\begin{equation*}
\mathcal{F}[X]=\int_{\Sigma} \gamma(\nu) d \Sigma \tag{1}
\end{equation*}
$$

where $d \Sigma$ is the area element of $X$. Such a functional is used to model anisotropic surface energies. Applications can be found in many branches of the physical sciences including metallurgy and crystallography ([14, 15]). We will call $\mathcal{F}[X]$ the anisotropic energy of $X$, and $\gamma$ the energy density function.

We call stationary surfaces of (1) for compactlysupported variations $\gamma$-minimal surfaces. It is obvious that, for $\gamma \equiv 1$, $\gamma$-minimal surfaces are usual minimal surfaces.

Denote by $D \gamma$ and $D^{2} \gamma$ the gradient and the Hessian of $\gamma$ on $\Omega$, respectively. Denote by 1 the identity endomorphism field on the tangent space $T_{\nu}\left(S^{2}\right)$. If the matrix $D^{2} \gamma+\gamma 1$ is non-singular at each point $\nu$ in $\Omega$, a mapping $Y: \Omega \rightarrow \mathbf{R}^{3}$ defined by $Y(\nu)=D \gamma+\gamma(\nu) \nu$ is an immersion and $Y$ defines the uniquely determined immersed surface with unit normal $\nu$ whose support function coincides with $\gamma$, that is $\gamma(\nu)=\langle Y(\nu), \nu\rangle$ holds. We say that $Y$ is the standard body for $\gamma$. (As for the terminology "standard body", we quote
[12].) We will sometimes use the symbol $M_{\gamma}$ to represent the mapping $Y$ or the image $Y(\Omega)$ of $Y$.
We say that $\gamma: \Omega \rightarrow \mathbf{R}_{+}$satisfies the convexity condition, if the matrix $D^{2} \gamma+\gamma 1$ is positive definite at each point $\nu$ in $\Omega$. In this case, the standard body $M_{\gamma}$ for $\gamma$ is strongly convex (that is, the principal curvatures of $M_{\gamma}$ are positive everywhere), and the functional $\mathcal{F}$ appearing in (1) is called a constant coefficient parametric elliptic functional, and stationary surfaces are extensively studied in recent years.

In this paper, we do not assume the convexity condition. By this generalization, we obtain a more variety of important examples. For example, zero mean curvature immersions in the Lorentz-Minkowski space $\mathbf{R}_{1}^{3}:=\left\{\left(x_{1}, x_{2}, x_{3}\right) \in\right.$ $\left.\mathbf{R}^{3} ; d s^{2}=d x_{1}^{2}+d x_{2}^{2}-d x_{3}^{2}\right\}$ arise as $\gamma$-minimal surfaces for a certain simple function $\gamma$ as follows (cf. $\S 3$ ).
Theorem 1. Set $\Omega_{1}:=\left\{\nu=\left(\nu_{1}, \nu_{2}, \nu_{3}\right) \in S^{2} ;\left|\nu_{3}\right|>\right.$ $\sqrt{2} / 2\}, \Omega_{2}:=\left\{\nu \in S^{2} ;\left|\nu_{3}\right|<\sqrt{2} / 2\right\}$. Define a function $\gamma: S^{2} \rightarrow \mathbf{R}$ as $\gamma(\nu)=\sqrt{\left|\nu_{3}^{2}-\nu_{1}^{2}-\nu_{2}^{2}\right|}=\sqrt{\left|2 \nu_{3}^{2}-1\right|}$. Then, an immersion $X: \Sigma \rightarrow \mathbf{R}^{3}$ with Gauss image $\nu(\Sigma) \subset$ $\Omega_{1} \cup \Omega_{2}$ is $\gamma$-minimal if and only if the mean curvature of $X$ is zero as an immersed surface in $\mathbf{R}_{1}^{3}$.
This result indicates that the recent investigations about zero mean curvature surfaces in $\mathbf{R}_{1}^{3}$ changing their causal type across null curves (regular curves whose velocity vector fields are lightlike) or lightlike lines from spacelike zero mean curvature surfaces to timelike zero mean curvature surfaces $([3,6,5,4,2])$ should be very natural and reasonable. Probably the most well-known example of such surfaces is the right helicoid with the timelike axis as its axis, which changes its causal type across a null curve from a spacelike zero mean curvature surface to a timelike zero mean curvature surface $([3,6])$. In $\S 4$, we will show a more
general remarkable result as follows.
Theorem 2. Let $\gamma: \Omega \rightarrow \mathbf{R}_{+}$be a positive $C^{\infty}$ function on a nonempty open set $\Omega$ in $S^{2}$. Assume that the matrix $D^{2} \gamma+\gamma 1$ is non-singular at each point $\nu$ in $\Omega$. Assume also that $\gamma$ is axisymmetric and not a constant function. Let $X: \Sigma \rightarrow \mathbf{R}^{3}$ be an immersion which is compatible with $\gamma$. Then, $X$ is both minimal and $\gamma$-minimal if and only if it is a part of either a plane or a right helicoid whose axis is parallel to the axis of $\gamma$.

This result is a generalization of [7, Theorem 4.2] and a refinement of [9, Proposition III.1]. [7, Theorem 4.2] proves that a spacelike plane and the spacelike part of a right helicoid whose axis is parallel to the timelike axis are only both a minimal surface in the euclidean space $\mathbf{R}^{3}$ and a spacelike zero mean curvature surface in $\mathbf{R}_{1}^{3}$. [9, Proposition III.1] proves that a right helicoid is a $\gamma$-minimal surface for any axisymmetric $\gamma$ whose axis is parallel to the axis of the helicoid itself.

Theorem 2 combined with Theorem 1 implies the following:
Corollary 1. A spacelike plane and the spacelike part of a right helicoid whose axis is parallel to the $x_{3}$-axis are only both a minimal surface in the euclidean space $\mathbf{R}^{3}$ and $a$ spacelike zero mean curvature surface in $\mathbf{R}_{1}^{3}$. Also, a timelike plane and the timelike part of a right helicoid whose axis is parallel to the $x_{3}$-axis are only both a minimal surface in $\mathbf{R}^{3}$ and a timelike zero mean curvature surface in $\mathbf{R}_{1}^{3}$.

In general, it is not easy to construct examples of $\gamma$ minimal surfaces. For any axisymmetric energy density function $\gamma$, there exist $\gamma$-minimal surfaces which are also symmetric with respect to the same axis as $\gamma$. The existence theorem and a certain kind of representation formula of these surfaces were given in [8] and they were called anisotropic catenoid. Although the convexity condition for $\gamma$ was assumed in [8], the method there works also for non convex $\gamma$. In this paper, for certain classes of $\gamma$, we will give another type of examples of $\gamma$-minimal surfaces which are foliated by parallel circles but are not surfaces of revolution. We will call them $\gamma$-minimal surfaces of Riemann-type after Riemann's minimal surfaces in $\mathbf{R}^{3}$.
Proposition 1. Let $\gamma: \Omega \rightarrow \mathbf{R}_{+}$be a positive $C^{\infty}$ function on a nonempty open set $\Omega$ in $S^{2}$. Assume that the matrix $D^{2} \gamma+\gamma 1$ is non-singular at each point $\nu \in \Omega$. We also assume that the standard body $M_{\gamma}$ for $\gamma$ is a quadric surface of revolution. Then, there are $\gamma$-minimal surfaces of Riemann-type.

From Theorem 1, we see that spacelike and timelike zero mean curvature surfaces of Riemann-type in $\mathbf{R}_{1}^{3}$ are obtained as special cases of surfaces given by Proposition 1. Actually, for $\left.\gamma\right|_{\Omega_{1}}$ in Theorem $1, M_{\gamma}$ is a hyperboloid of two sheets, and for $\left.\gamma\right|_{\Omega_{2}}, M_{\gamma}$ is a hyperboloid of one sheet ( $\S 5$, Lemma 5).

We should remark that zero mean curvature surfaces of Riemann-type in $\mathbf{R}_{1}^{3}$ were studied also in [10, 11].

In $\S 5$, for $\gamma$ satisfying the assumption in Proposition 1, we will give explicit parameter representations of all $\gamma$ -
minimal surfaces foliated by circles contained in parallel planes which are orthogonal to the rotation axis of $M_{\gamma}$ (Proposition 3). Actually, Proposition 1 is a corollary of Proposition 3.

Some of the results in this article can be generalized to hypersurfaces in $\mathbf{R}^{n+1}$.

## 2. Preliminaries

In this section, we give the definitions of the Wulff shape, anisotropic mean curvature, and their fundamental properties and representation formulas. We quote $[12,1,8]$ as references.

Let $\gamma: \Omega \rightarrow \mathbf{R}_{+}$be a positive $C^{\infty}$ function on a nonempty open set $\Omega$ of the unit sphere $S^{2}$. Assume that the matrix $D^{2} \gamma+\gamma 1$ is non-singular at each point $\nu$ in $\Omega$.

If $\Omega=S^{2}$, then, for any $V>0$, there exists a uniquely determined (up to translations in $\mathbf{R}^{3}$ ) convex surface $W(V)$ such that $W(V)$ attains the minimum of $\mathcal{F}$ among all closed piecewise smooth surfaces in $\mathbf{R}^{3}$ enclosing the 3dimensional volume $V([13])$. For the special value $V_{0}:=$ $(1 / 3) \int_{S^{2}} \gamma(\nu) d S^{2}, W\left(V_{0}\right)$ is called the Wulff shape for $\gamma$, and we will denote it by $W$. In the special case where $\gamma \equiv 1, \mathcal{F}[X]$ is the usual area of the surface $X$ and $W$ is the unit sphere $S^{2}$. In general, $W$ is not smooth. $W$ is a smooth strongly convex surface if and only if $\gamma$ satisfies the convexity condition (see $\S 1$ ). In this case, $W$ can be parametrized by the smooth mapping

$$
Y: S^{2} \rightarrow \mathbf{R}^{3}, \quad Y(\nu)=D \gamma+\gamma(\nu) \nu
$$

where we regard $D \gamma$ at $\nu \in S^{2}$ as a point in $\mathbf{R}^{3}$ in the canonical manner. We remark that the outward unit normal to $W$ at point $Y(\nu)$ coincides with $\nu$. And the function $\gamma$ coincides with the support function of $W$, that is $\gamma(\nu)=\langle Y(\nu), \nu\rangle$, where $\langle$,$\rangle is the inner product in \mathbf{R}^{3}$. This means that $W$ is the standard body for $\gamma$.

Let $X: \Sigma \rightarrow \mathbf{R}^{3}$ be an immersion. By parallel translation in $\mathbf{R}^{3}, D \gamma$ may be considered as a smooth tangent vector field along $X$. Let $X_{\epsilon}=X+\epsilon \delta X+\mathcal{O}\left(\epsilon^{2}\right)$ be a smooth, compactly supported variation of $X$. The anisotropic mean curvature $\Lambda$ of $X$ is defined by the first variation formula ([8])

$$
\begin{gather*}
\delta \mathcal{F}:=\partial_{\epsilon} \mathcal{F}\left[X_{\epsilon}\right]_{\epsilon=0}=-\int_{\Sigma} \Lambda\langle\delta X, \nu\rangle d \Sigma  \tag{2}\\
\Lambda:=-\operatorname{trace}_{\Sigma}\left(D^{2} \gamma+\gamma 1\right) d \nu=-\operatorname{div}_{\Sigma} D \gamma+2 H \gamma \tag{3}
\end{gather*}
$$

where $H$ is the mean curvature of $X$. Hence, $\gamma$-minimal surfaces are immersed surfaces whose anisotropic mean curvature $\Lambda$ vanishes at every point. Since the first variation of the "enclosed volume" $V[X]:=(1 / 3) \int_{\Sigma}\langle X, \nu\rangle d \Sigma$ satisfies

$$
\delta V[X]=\int_{\Sigma}\langle\delta X, \nu\rangle d \Sigma
$$

the equation $\Lambda \equiv$ constant characterizes critical points of $\mathcal{F}$ with the enclosed volume constrained to be a constant. If $\Lambda$ is constant, $X$ is called a surface of constant anisotropic mean curvature. In the case where $\gamma \equiv 1, \Lambda=2 H$ holds.

Now we extend the function $\gamma$ in a homogeneous way to a function $\tilde{\gamma}$ as follows.
(i) $\tilde{\gamma}(X)=0$ if and only if $X=\mathbf{0}$.
(ii) positive homogeneity of degree one:

$$
\tilde{\gamma}(r X)=r \gamma(X), \quad \forall r \geq 0, X \in \Omega
$$

In the special case where $\gamma(X) \equiv 1, \tilde{\gamma}(X) \equiv|X|$.
Let us consider a surface which is a graph of a $C^{\infty}$ function $\varphi: \Sigma\left(\subset \mathbf{R}^{2}\right) \rightarrow \mathbf{R}$ as follows:

$$
X: \Sigma \rightarrow \mathbf{R}^{3}, \quad X\left(x_{1}, x_{2}\right)=\left(x_{1}, x_{2}, \varphi\left(x_{1}, x_{2}\right)\right)
$$

The unit normal $\nu=\left(\nu_{1}, \nu_{2}, \nu_{3}\right)$ to $X$ is given by

$$
\begin{equation*}
\nu=\frac{\left(-\varphi_{1},-\varphi_{2}, 1\right)}{\left(1+|D \varphi|^{2}\right)^{1 / 2}}, \tag{4}
\end{equation*}
$$

where

$$
\varphi_{1}:=\varphi_{x_{1}}, \quad \varphi_{2}:=\varphi_{x_{2}}, \quad D \varphi:=\left(\varphi_{1}, \varphi_{2}\right)
$$

Lemma 1. Set $\varphi_{i j}:=\varphi_{x_{i} x_{j}}$ for $i, j=1,2$. Then

$$
\begin{equation*}
\Lambda=\left.\sum_{i, j=1,2} \tilde{\gamma}_{x_{i} x_{j}}\right|_{X=(-D \varphi, 1)} \varphi_{i j} \tag{5}
\end{equation*}
$$

holds. In the special case where $\tilde{\gamma}(X) \equiv|X|$, the right hand side of (5) is

$$
\begin{equation*}
\frac{\left(1+\varphi_{2}^{2}\right) \varphi_{11}-2 \varphi_{1} \varphi_{2} \varphi_{12}+\left(1+\varphi_{1}^{2}\right) \varphi_{22}}{\left(\varphi_{1}^{2}+\varphi_{2}^{2}+1\right)^{3 / 2}} \tag{6}
\end{equation*}
$$

which is the twice of the mean curvature $H$ of $X$.
Proof. In the integrals below, we will write $\varphi\left(u_{1}, u_{2}\right)$, $\left(\left(u_{1}, u_{2}\right) \in \Sigma\right)$, in order to avoid confusion. We have

$$
\begin{aligned}
\mathcal{F}[X] & =\iint_{\Sigma} \gamma(\nu)\left(1+\varphi_{1}^{2}+\varphi_{2}^{2}\right)^{1 / 2} d u_{1} d u_{2} \\
& =\iint_{\Sigma} \tilde{\gamma}\left(\left(-\varphi_{1},-\varphi_{2}, 1\right)\right) d u_{1} d u_{2}
\end{aligned}
$$

Let $X_{\epsilon}=\left(x_{1}, x_{2}, \varphi\left(\epsilon, x_{1}, x_{2}\right)\right)$ be an arbitrary compactlysupported variation of $X$. We will compute the first variation of $\mathcal{F}$. We may suppose that $\Sigma$ is the support of $X_{\epsilon}$ and $\left.X_{\epsilon}\right|_{\partial \Sigma}=\left.X\right|_{\partial \Sigma}$ holds. We compute

$$
\begin{aligned}
\delta \mathcal{F}= & \iint_{\Sigma}\left(\tilde{\gamma}\left(\left(-\varphi_{1},-\varphi_{2}, 1\right)\right)\right)_{\epsilon} d u_{1} d u_{2} \\
= & \iint_{\Sigma} \tilde{\gamma}_{x_{1}} \cdot\left(-\varphi_{1 \epsilon}\right)+\tilde{\gamma}_{x_{2}} \cdot\left(-\varphi_{2 \epsilon}\right) d u_{1} d u_{2} \\
= & \iint_{\Sigma} \frac{\left.\partial \tilde{\gamma}_{x_{1}}\right|_{(-D \varphi, 1)}}{\partial u_{1}} \varphi_{\epsilon}+\frac{\left.\partial \tilde{\gamma}_{x_{2}}\right|_{(-D \varphi, 1)}}{\partial u_{2}} \varphi_{\epsilon} d u_{1} d u_{2} \\
& \quad-\iint_{\Sigma}\left(\tilde{\gamma}_{x_{1}} \varphi_{\epsilon}\right)_{u_{1}}+\left(\tilde{\gamma}_{x_{2}} \varphi_{\epsilon}\right)_{u_{2}} d u_{1} d u_{2}
\end{aligned}
$$

By the partial differentiation, the last term of the above equation becomes

$$
\int_{\partial \Sigma}\left(-\tilde{\gamma}_{x_{2}} \varphi_{\epsilon} d u_{1}+\tilde{\gamma}_{x_{1}} \varphi_{\epsilon} d u_{2}\right)=0
$$

because $\varphi_{\epsilon}=0$ on $\partial \Sigma$. Therefore,

$$
\begin{align*}
\delta \mathcal{F} & =-\iint_{\Sigma}\left(\tilde{\gamma}_{x_{1} x_{1}} \varphi_{11}+2 \tilde{\gamma}_{x_{1} x_{2}} \varphi_{12}+\tilde{\gamma}_{x_{2} x_{2}} \varphi_{22}\right) \varphi_{\epsilon} d u_{1} d u_{2} \\
& =-\iint_{\Sigma}\left(\left.\sum_{i, j=1,2} \tilde{\gamma}_{x_{i} x_{j}}\right|_{X=(-D \varphi, 1)} \varphi_{i j}\right)\langle\delta X, \nu\rangle d \Sigma, \tag{7}
\end{align*}
$$

here we used (4) and the followings:

$$
\delta X=\left(0,0, \varphi_{\epsilon}\right), \quad d \Sigma=\left(1+|D \varphi|^{2}\right)^{1 / 2} d u_{1} d u_{2}
$$

In view of (2), (7) implies (5). By a direct computation, we obtain (6).

We will give another representation of the anisotropic mean curvature. Let $X: \Sigma \rightarrow \mathbf{R}^{3}$ be an immersion with Gauss map $\nu$. Let $\left\{e_{1}, e_{2}\right\}$ be a locally defined frame on $S^{2}$ such that $\left(D^{2} \gamma+\gamma 1\right) e_{i}=\left(1 / \mu_{i}\right) e_{i}$. Note that the basis $\left\{e_{1}, e_{2}\right\}$ at $\nu(p)$ also serves as an orthogonal basis for the tangent plane of $X$ at $p$. Let $\left(-w_{i j}\right)$ be the matrix representing $d \nu$ with respect to this basis. Then

$$
\left(D^{2} \gamma+\gamma 1\right) d \nu=\left(\begin{array}{ll}
-w_{11} / \mu_{1} & -w_{12} / \mu_{1} \\
-w_{21} / \mu_{2} & -w_{22} / \mu_{2}
\end{array}\right)
$$

This with (3) gives

$$
\begin{equation*}
\Lambda=w_{11} / \mu_{1}+w_{22} / \mu_{2} \tag{8}
\end{equation*}
$$

Note that $D^{2} \gamma+\gamma 1$ is the inverse of the differential of the Gauss map of $M_{\gamma}$ and so its eigenvalues $1 / \mu_{j}$ are the negatives of the reciprocals of the principal curvatures of the standard body $M_{\gamma}$ with respect to the outward unit normal.

For an axisymmetric $\gamma, \mu_{i}$ 's are represented in terms of $\gamma$ as follows:
Lemma 2. Let $\gamma: \Omega \rightarrow \mathbf{R}_{+}$be a positive $C^{\infty}$ function on a nonempty open set $\Omega$ of the unit sphere $S^{2}$. Assume that the matrix $D^{2} \gamma+\gamma 1$ is non-singular at each point $\nu$ in $\Omega$. Assume also that $\gamma$ is axisymmetric, say $\gamma(\nu)=$ $\gamma\left(\nu_{3}\right)$. Then the standard body $M_{\gamma}$ for $\gamma$ is also symmetric with respect to the $x_{3}$-axis. Denote by $\mu_{1}, \mu_{2}$ the principal curvatures of $M_{\gamma}$ with respect to the normal $-\nu$. We let $\mu_{1}$ be the curvature of the generating curve of $M_{\gamma}$. Then

$$
\begin{equation*}
\mu_{1}^{-1}=\left(1-\nu_{3}^{2}\right) \gamma^{\prime \prime}+\mu_{2}^{-1}, \quad \mu_{2}^{-1}=\gamma-\nu_{3} \gamma^{\prime} \tag{9}
\end{equation*}
$$

holds.
Proof. The proof is the same as the proof of the same formulas for the case where $\gamma$ satisfies the convexity condition which was given in [8, Section 5].

## 3. Proof of Theorem 1

In this section, we give a proof of Theorem 1 which was given in the introduction.

Denote by $\langle,\rangle_{L}$ the scalar product for the Minkowski metric $d x_{1}^{2}+d x_{2}^{2}-d x_{3}^{2}$ in $\mathbf{R}_{1}^{3}$. Let $X: \Sigma\left(\subset \mathbf{R}^{2}\right) \rightarrow \mathbf{R}_{1}^{3}$ be a spacelike or timelike immersed surface. Let $\left(u_{1}, u_{2}\right)$ be
local coordinates of $\Sigma$. Denote by $H_{L}$ the mean curvature of $X$. That is, $H_{L}$ is defined by

$$
H_{L}=\frac{\tilde{h}_{11} \tilde{g}_{22}-2 \tilde{h}_{12} \tilde{g}_{12}+\tilde{h}_{22} \tilde{g}_{11}}{2\left(\tilde{g}_{11} \tilde{g}_{22}-\tilde{g}_{12}^{2}\right)}
$$

where $\tilde{g}_{i j}:=\left\langle X_{u_{i}}, X_{u_{j}}\right\rangle_{L}, \tilde{h}_{i j}:=\left\langle X_{u_{i} u_{j}}, \nu^{L}\right\rangle_{L}$ for $i, j=$ 1,2 , and $\nu^{L}$ is the unit normal vector field along $X$ for the Minkowski metric. Let $A_{L}[X]$ be the area of $X$ defined by

$$
A_{L}[X]:=\int_{\Sigma} d \Sigma_{L}, \quad\left(d \Sigma_{L}:=\left|\operatorname{det}\left(\tilde{g}_{i j}\right)\right| d u_{1} d u_{2}\right)
$$

Let $X_{\epsilon}$ be an arbitrary compactly-supported variation of $X$. We will compute the first variation of $A_{L}$. We may suppose that $\Sigma$ is the support of $X_{\epsilon}$ and $\left.X_{\epsilon}\right|_{\partial \Sigma}=\left.X\right|_{\partial \Sigma}$ holds. Set the variation vector field as

$$
\delta X:=\partial_{\epsilon}\left(X_{\epsilon}\right)_{\epsilon=0}=\xi+f \nu^{L}, \quad\left(\xi=\sum_{i=1,2} \xi^{i} X_{u_{i}}\right)
$$

Then we have the following.
Proposition 2. In the above setting, it holds that

$$
\partial_{\epsilon} A_{L}\left[X_{\epsilon}\right]_{\epsilon=0}=-2 \int_{\Sigma} f H_{L} d \Sigma_{L}
$$

Proof. We here give a proof in the case where $X$ is timelike. By a similar argument, we can prove this in the case where $X$ is spacelike. We have

$$
A_{L}\left[X_{\epsilon}\right]=\int_{\Sigma} \sqrt{-\tilde{g}_{11}^{\epsilon} \tilde{g}_{22}^{\epsilon}+\left(\tilde{g}_{12}^{\epsilon}\right)^{2}} d u_{1} d u_{2}
$$

where $\tilde{g}_{i j}^{\epsilon}=\left\langle\left(X_{\epsilon}\right)_{u_{i}},\left(X_{\epsilon}\right)_{u_{j}}\right\rangle_{L}$ for $i, j=1,2$. Then,

$$
\begin{aligned}
\partial_{\varepsilon} A_{L}\left[X_{\epsilon}\right] & =\int_{\Sigma} \partial_{\varepsilon}\left(\sqrt{-\tilde{g}_{11}^{\epsilon} \tilde{g}_{22}^{\epsilon}+\left(\tilde{g}_{12}^{\epsilon}\right)^{2}}\right) d u_{1} d u_{2} \\
& =\int_{\Sigma} \frac{\partial_{\varepsilon}\left(-\tilde{g}_{11}^{\epsilon} \tilde{g}_{22}^{\epsilon}+\left(\tilde{g}_{12}^{\epsilon}\right)^{2}\right)}{2 \sqrt{-\tilde{g}_{11}^{\epsilon} \tilde{g}_{22}^{\epsilon}+\left(\tilde{g}_{12}^{\epsilon}\right)^{2}}} d u_{1} d u_{2} \\
& =\int_{\Sigma} \frac{\partial_{\varepsilon}\left(\tilde{g}_{11}^{\epsilon} \tilde{g}_{22}^{\epsilon}-\left(\tilde{g}_{12}^{\epsilon}\right)^{2}\right)}{2\left(\tilde{g}_{11}^{\epsilon} \tilde{g}_{22}^{\epsilon}-\left(\tilde{g}_{12}^{\epsilon}\right)^{2}\right)} d \Sigma_{L} \\
& =\int_{\Sigma} \frac{\tilde{g}_{22}^{\epsilon} \partial_{\varepsilon} \tilde{g}_{11}^{\epsilon}+\tilde{g}_{11}^{\epsilon} \partial_{\varepsilon} \tilde{g}_{22}^{\epsilon}-2 \tilde{g}_{12}^{\epsilon} \partial_{\varepsilon} \tilde{g}_{12}^{\epsilon}}{2\left(\tilde{g}_{11}^{\epsilon} \tilde{g}_{22}^{\epsilon}-\left(\tilde{g}_{12}^{\epsilon}\right)^{2}\right)} d \Sigma_{L}
\end{aligned}
$$

holds. By a direct calculation, we have

$$
\partial_{\varepsilon}\left(\tilde{g}_{i j}^{\epsilon}\right)_{\epsilon=0}=\left\langle\xi_{u_{i}}, X_{u_{j}}\right\rangle_{L}+\left\langle X_{u_{i}}, \xi_{u_{j}}\right\rangle_{L}-2 f \tilde{h}_{i j}
$$

for $i, j=1,2$. Applying the divergence theorem, it follows that

$$
\begin{aligned}
\partial_{\varepsilon} A_{L}\left[X_{\epsilon}\right]_{\epsilon=0} & =\int_{\Sigma} \sum_{i=1,2}\left(\tilde{g}^{i j}\left\langle\xi_{u_{i}}, X_{u_{j}}\right\rangle_{L}-f \tilde{g}^{i j} \tilde{h}_{i j}\right) d \Sigma_{L} \\
& =\int_{\Sigma}\left(\operatorname{div} \xi-2 f H_{L}\right) d \Sigma_{L}=-2 \int_{\Sigma} f H_{L} d \Sigma_{L}
\end{aligned}
$$

where we denote by $\left(\tilde{g}^{i j}\right)$ the inverse matrix of $\left(\tilde{g}_{i j}\right)$.

Proof of Theorem 1. First we assume that the surface is a graph of a $C^{\infty}$ function $\varphi: \Sigma\left(\subset \mathbf{R}^{2}\right) \rightarrow \mathbf{R}$ as follows:

$$
\begin{equation*}
X: \Sigma \rightarrow \mathbf{R}^{3}, \quad X\left(x_{1}, x_{2}\right)=\left(x_{1}, x_{2}, \varphi\left(x_{1}, x_{2}\right)\right) \tag{10}
\end{equation*}
$$

The area element $d \Sigma_{L}$ of $X$ is given by

$$
d \Sigma_{L}=\left|1-\varphi_{1}^{2}-\varphi_{2}^{2}\right|^{1 / 2} d x_{1} d x_{2}
$$

On the other hand, the unit normal $\nu=\left(\nu_{1}, \nu_{2}, \nu_{3}\right)$ to $X$ and the area element $d \Sigma$ of $X$ for the euclidean metric are

$$
\nu=\frac{\left(-\varphi_{1},-\varphi_{2}, 1\right)}{\left(1+\varphi_{1}^{2}+\varphi_{2}^{2}\right)^{1 / 2}}, \quad d \Sigma=\left(1+\varphi_{1}^{2}+\varphi_{2}^{2}\right)^{1 / 2} d u_{1} d u_{2}
$$

Hence,

$$
d \Sigma_{L}=\left(\frac{\left|1-\varphi_{1}^{2}-\varphi_{2}^{2}\right|}{1+\varphi_{1}^{2}+\varphi_{2}^{2}}\right)^{1 / 2} d \Sigma=\left|\nu_{3}^{2}-\nu_{1}^{2}-\nu_{2}^{2}\right|^{1 / 2} d \Sigma
$$

Therefore, by (2) and Proposition $2, \Lambda \equiv 0$ if and only if $H_{L} \equiv 0$.

Next, we consider the case where the considered surface $\Sigma$ cannot be represented as a graph like (10). It is sufficient to consider the case where the image of the Gauss map of $\Sigma$ is contained in the equator $\left\{\left(x_{1}, x_{2}, 0\right) \in S^{2}\right\}$. In this case, $\Sigma$ is timelike. It is proved that $\Sigma$ is represented as

$$
X(s, t)=\left(x_{1}(s), x_{2}(s), t\right)
$$

where $C(s):=\left(x_{1}(s), x_{2}(s)\right)$ is a smooth plane curve. Denote by $\kappa$ the curvature of $C$. Note that $\gamma$ can be represented as $\gamma(\nu)=\gamma\left(\nu_{3}\right)$. Then from (8), $\Lambda(s, t)=\gamma(0) \kappa(s)$ holds. On the other hand, $H_{L}=\kappa / 2$ holds. Hence, $\Lambda \equiv 0$ if and only if $H_{L} \equiv 0$.

## 4. Proof of Theorem 2

Let $\left(x_{1}, x_{2}, x_{3}\right)$ be the standard coordinates in $\mathbf{R}^{3}$. We assume that $\gamma$ is symmetric with respect to the $x_{3}$-axis without loss of generality. So we can write $\gamma=\gamma\left(\nu_{3}\right)$. Assume that $\gamma$ is not a constant function.

Denote by $\Sigma$ the considered surface. First assume that $\Sigma$ is represented as $x_{3}=\varphi\left(x_{1}, x_{2}\right)$. As in $\S 2$, we will write

$$
\varphi_{i}:=\varphi_{x_{i}}, \quad \varphi_{i j}:=\varphi_{x_{i} x_{j}}, \quad(i, j=1,2)
$$

By the formula (5) and a simple but long computation, we have

$$
\begin{aligned}
\Lambda=2 H\left(\gamma-\frac{\gamma^{\prime}}{\left(1+\varphi_{1}^{2}+\varphi_{2}^{2}\right)^{1 / 2}}-\right. & \left.\frac{\gamma^{\prime \prime}}{1+\varphi_{1}^{2}+\varphi_{2}^{2}}\right) \\
& +\frac{\gamma^{\prime \prime}\left(\varphi_{11}+\varphi_{22}\right)}{\left(1+\varphi_{1}^{2}+\varphi_{2}^{2}\right)^{3 / 2}}
\end{aligned}
$$

Hence, if $\Lambda=H=0$ holds, then $\gamma^{\prime \prime}\left(\varphi_{11}+\varphi_{22}\right)=0$ holds. Since

$$
0=H=\frac{\left(1+\varphi_{2}^{2}\right) \varphi_{11}-2 \varphi_{1} \varphi_{2} \varphi_{12}+\left(1+\varphi_{1}^{2}\right) \varphi_{22}}{2\left(\varphi_{1}^{2}+\varphi_{2}^{2}+1\right)^{3 / 2}}
$$

we obtain

$$
\begin{equation*}
\gamma^{\prime \prime}\left(\varphi_{2}^{2} \varphi_{11}-2 \varphi_{1} \varphi_{2} \varphi_{12}+\varphi_{1}^{2} \varphi_{22}\right)=0 \tag{11}
\end{equation*}
$$

Consider any contour line $\varphi\left(x_{1}(s), x_{2}(s)\right) \equiv$ constant, $(s$ is arc length of the curve $\left.C:\left(x_{1}(s), x_{2}(s)\right)\right)$, of $\Sigma$. Denote by $\kappa$ the curvature of $C$. Then,

## Lemma 3.

$$
\begin{equation*}
\left|\varphi_{2}^{2} \varphi_{11}-2 \varphi_{1} \varphi_{2} \varphi_{12}+\varphi_{1}^{2} \varphi_{22}\right|=|\kappa|\left(\varphi_{1}^{2}+\varphi_{2}^{2}\right)^{3 / 2} \tag{12}
\end{equation*}
$$

holds.
Proof. Denote by " $/$ " the derivative with respect to $s$. We differentiate $\varphi\left(x_{1}(s), x_{2}(s)\right) \equiv$ constant with respect to $s$ to obtain

$$
\begin{equation*}
\varphi_{1} x_{1}^{\prime}+\varphi_{2} x_{2}^{\prime}=0 \tag{13}
\end{equation*}
$$

Differentiate (13) again and use $\left(x_{1}^{\prime \prime}, x_{2}^{\prime \prime}\right)=\kappa\left(-x_{2}^{\prime}, x_{1}^{\prime}\right)$ to obtain

$$
\begin{equation*}
\varphi_{11}\left(x_{1}^{\prime}\right)^{2}+2 \varphi_{12} x_{1}^{\prime} x_{2}^{\prime}+\varphi_{22}\left(x_{2}^{\prime}\right)^{2}=\kappa\left(\varphi_{1} x_{2}^{\prime}-\varphi_{2} x_{1}^{\prime}\right) \tag{14}
\end{equation*}
$$

By using (13), (14), and the fact that $\left(x_{1}^{\prime}\right)^{2}+\left(x_{2}^{\prime}\right)^{2}=1$, we obtain (12).

Now we assume that the surface is not (a part of) a plane. We remark that it is sufficient to prove that the surface is a part of a right helicoid almost everywhere. So we assume that $\nu \neq(0,0, \pm 1)$ at any point in $\Sigma$, that is $\left(\varphi_{1}, \varphi_{2}\right)$ never coincides with $(0,0)$. Then, (11) combined with (12) shows that $\gamma^{\prime \prime} \equiv 0$ or $\kappa \equiv 0$ holds. If $\gamma^{\prime \prime} \equiv 0$, then, by Lemma $2, \mu_{1} \equiv \mu_{2}$ holds. This means that the standard body $M_{\gamma}$ for $\gamma$ is (a part of) a sphere, and hence $\gamma$ is a constant function, which contradicts the assumption. Hence $\kappa \equiv 0$ holds, and the curve $C$ is a straight line. Therefore, $\Sigma$ is a ruled surface. Because only planes and right helicoids are ruled surfaces which are minimal, $\Sigma$ is a right helicoid.

If $\Sigma$ is represented as $x_{3}=\varphi\left(x_{1}, x_{2}\right)$ in a connected neighborhood $U$ of a point $P_{0} \in \Sigma$, then, by the above argument, $U$ is a part of a right helicoid $M$. Since $\Sigma_{1}:=\Sigma \cap M$ is an open and closed subset of a connected set $\Sigma, \Sigma_{1}=\Sigma$ must hold. This means that $\Sigma$ itself is a part of a right helicoid.

If $\Sigma$ is not represented as a graph $x_{3}=\varphi\left(x_{1}, x_{2}\right)$ at any point, then $\nu(P)$ is in the equator of $S^{2}$ for any $P \in \Sigma$. Hence the Gauss curvature $K$ of $\Sigma$ vanishes at any point. Since $K \equiv 0 \equiv H, \Sigma$ is a plane which is parallel to the $x_{3}$-axis.

## 5. EXAMPLES

Let $\gamma: \Omega \rightarrow \mathbf{R}_{+}$be an axisymmetric positive $C^{\infty}$ function (say, $\gamma(\nu)=\gamma\left(\nu_{3}\right)$ ) on a nonempty open set $\Omega$ in $S^{2}$. Assume that the matrix $D^{2} \gamma+\gamma 1$ is non-singular at each point $\nu \in \Omega$.

In this section, we study a special type of cyclic surfaces, that is, surfaces foliated by circles in parallel planes which are orthogonal to the $x_{3}$-axis. So our surfaces are represented as follows:

$$
\begin{equation*}
X(\theta, t)=(r(t) \cos \theta+f(t), r(t) \sin \theta+g(t), t) \tag{15}
\end{equation*}
$$

As in Lemma 2, we denote by $\mu_{1}, \mu_{2}$ the principal curvatures of the standard body $M_{\gamma}$ with respect to the normal $-\nu$, here $\mu_{1}$ is the curvature of the generating curve of $M_{\gamma}$.
Lemma 4. The anisotropic mean curvature of $X$ in (15) is given by

$$
\begin{gather*}
\Lambda=\frac{r\left(r^{\prime \prime}+f^{\prime \prime} \cos \theta+g^{\prime \prime} \sin \theta\right)-\left(f^{\prime} \sin \theta-g^{\prime} \cos \theta\right)^{2}}{\mu_{1} r\left\{\left(r^{\prime}+f^{\prime} \cos \theta+g^{\prime} \sin \theta\right)^{2}+1\right\}^{\frac{3}{2}}} \\
-\frac{1}{\mu_{2} r \sqrt{\left(r^{\prime}+f^{\prime} \cos \theta+g^{\prime} \sin \theta\right)^{2}+1}} \tag{16}
\end{gather*}
$$

Proof. Let $\nu$ be the Gauss map of $X$ as usual. Let $\left\{e_{1}, e_{2}\right\}$ be a locally defined frame on $S^{2}$ such that $\left(D^{2} \gamma+\gamma 1\right) e_{i}=$ $\left(1 / \mu_{i}\right) e_{i}$. Note that the basis $\left\{e_{1}, e_{2}\right\}$ at $\nu(p)$ also serves as an orthogonal basis for the tangent plane of $X$ at $p$. As in $\S 2$, let $\left(-w_{i j}\right)$ be the matrix representing $d \nu$ with respect to this basis. Then

$$
\left(D^{2} \gamma+\gamma 1\right) d \nu=\left(\begin{array}{ll}
-w_{11} / \mu_{1} & -w_{12} / \mu_{1} \\
-w_{21} / \mu_{2} & -w_{22} / \mu_{2}
\end{array}\right)
$$

and

$$
\begin{equation*}
\Lambda=w_{11} / \mu_{1}+w_{22} / \mu_{2} \tag{17}
\end{equation*}
$$

holds. So, we will compute the matrix $\left(w_{i j}\right)$.
Let $\nu^{M}=\left(\nu_{1}^{M}, \nu_{2}^{M}, \nu_{3}^{M}\right)$ be the outward pointing unit normal to $M_{\gamma}$. Since $M_{\gamma}$ is a surface of revolution, $D^{2} \gamma+\gamma 1$ has eigendirections corresponding to

$$
\begin{equation*}
E_{1}=(0,0,1)-\nu_{3}^{M} \nu^{M}, \quad E_{2}=\nu^{M} \times E_{1} \tag{18}
\end{equation*}
$$

as long as the normal is not vertical. $E_{1}, E_{2}$ define an orthonormal basis $\left\{e_{1}, e_{2}\right\}$ on $T S^{2}$ as long as $X$ does not intersect with the vertical axis.

Set $g_{11}=\left\langle X_{\theta}, X_{\theta}\right\rangle, g_{12}=g_{21}=\left\langle X_{\theta}, X_{t}\right\rangle, g_{22}=$ $\left\langle X_{t}, X_{t}\right\rangle, h_{11}=\left\langle X_{\theta \theta}, \nu\right\rangle, h_{12}=h_{21}=\left\langle X_{\theta t}, \nu\right\rangle, h_{22}=$ $\left\langle X_{t t}, \nu\right\rangle$. And set

$$
\Delta:=r^{\prime}+f^{\prime} \cos \theta+g^{\prime} \sin \theta
$$

Then,

$$
\begin{aligned}
& g_{11}=r^{2}, \\
& g_{12}=-r f^{\prime} \sin \theta+r g^{\prime} \cos \theta, \\
& g_{22}=\left(r^{\prime}\right)^{2}+2 r^{\prime} f^{\prime} \cos \theta+2 r^{\prime} g^{\prime} \sin \theta+\left(f^{\prime}\right)^{2}+\left(g^{\prime}\right)^{2}+1, \\
& h_{11}=\frac{-r}{\sqrt{\Delta^{2}+1}}, \\
& h_{12}=0 \\
& h_{22}=\frac{r^{\prime \prime}+f^{\prime \prime} \cos \theta+g^{\prime \prime} \sin \theta}{\sqrt{\Delta^{2}+1}},
\end{aligned}
$$

and
$\nu=\left(\nu_{1}, \nu_{2}, \nu_{3}\right):=\frac{X_{\theta} \times X_{t}}{\left|X_{\theta} \times X_{t}\right|}=\frac{1}{\sqrt{\Delta^{2}+1}}(\cos \theta, \sin \theta,-\Delta)$.
We have

$$
\tilde{E}_{1}:=(0,0,1)-\nu_{3} \nu=\frac{1}{\Delta^{2}+1}(\Delta \cos \theta, \Delta \sin \theta, 1) .
$$

Hence,

$$
\begin{align*}
& e_{1}=\frac{\tilde{E}_{1}}{\left|\tilde{E}_{1}\right|}=\frac{1}{\sqrt{\Delta^{2}+1}}(\Delta \cos \theta, \Delta \sin \theta, 1),  \tag{19}\\
& e_{2}=\nu \times e_{1}=(\sin \theta,-\cos \theta, 0) . \tag{20}
\end{align*}
$$

Now we take a coordinate transformation $\theta(u, v), t(u, v)$ so that, at an arbitrary fixed point $\left(u_{0}, v_{0}\right)$,

$$
\frac{\partial X}{\partial u}=e_{1}, \quad \frac{\partial X}{\partial v}=e_{2}
$$

are satisfied. Then, we have

$$
\begin{equation*}
\frac{\partial X}{\partial \theta} \theta_{u}+\frac{\partial X}{\partial t} t_{u}=e_{1}, \quad \frac{\partial X}{\partial \theta} \theta_{v}+\frac{\partial X}{\partial t} t_{v}=e_{2} \tag{21}
\end{equation*}
$$

Inserting $X_{\theta}=(-r \sin \theta, r \cos \theta, 0), X_{t}=\left(r^{\prime} \cos \theta+\right.$ $\left.f^{\prime}, r^{\prime} \sin \theta+g^{\prime}, 1\right)$, (19), and (20) to (21), we obtain

$$
J:=\left(\begin{array}{ll}
\theta_{u} & \theta_{v} \\
t_{u} & t_{v}
\end{array}\right)=\left(\begin{array}{cc}
\frac{f^{\prime} \sin \theta-g^{\prime} \cos \theta}{r \sqrt{\Delta^{2}+1}} & -\frac{1}{r} \\
\frac{1}{\sqrt{\Delta^{2}+1}} & 0
\end{array}\right), \quad \operatorname{det} J>0 .
$$

Let $\left(w_{i j}\right),\left(\tilde{w}_{i j}\right)$ be the Weingarten mappings for $X(u, v)$, $X(\theta, t)$, respectively. Then,

$$
\begin{gathered}
\left(\tilde{w}_{i j}\right)=\left(g_{i j}\right)^{-1}\left(h_{i j}\right) \\
\left(w_{i j}\right)=\left(\begin{array}{cc}
\theta_{u} & \theta_{v} \\
t_{u} & t_{v}
\end{array}\right)^{-1}\left(\tilde{w}_{i j}\right)\left(\begin{array}{ll}
\theta_{u} & \theta_{v} \\
t_{u} & t_{v}
\end{array}\right)=J^{-1}\left(\tilde{w}_{i j}\right) J
\end{gathered}
$$

Hence, by a computation, we obtain

$$
\begin{aligned}
& w_{11}=\frac{r\left(r^{\prime \prime}+f^{\prime \prime} \cos \theta+g^{\prime \prime} \sin \theta\right)-\left(f^{\prime} \sin \theta-g^{\prime} \cos \theta\right)^{2}}{r\left(\Delta^{2}+1\right)^{\frac{3}{2}}} \\
& w_{22}=-\frac{1}{r \sqrt{\Delta^{2}+1}}
\end{aligned}
$$

This with (17) gives (16).
Now we assume that the standard body $M_{\gamma}$ for $\gamma$ is a quadric surface of revolution. Then, by homothety and translation, $M_{\gamma}$ is one of the followings:
(I) a spheroid: $x_{1}^{2}+x_{2}^{2}+\frac{x_{3}^{2}}{a^{2}}=1$,
(II) a hyperboloid of two sheets: $x_{1}^{2}+x_{2}^{2}-\frac{x_{3}^{2}}{a^{2}}=-1$,
(III) a hyperboloid of one sheet: $x_{1}^{2}+x_{2}^{2}-\frac{x_{3}^{2}}{a^{2}}=1$,
(IV) a circular paraboloid: $x_{3}=a\left(x_{1}^{2}+x_{2}^{2}\right)$,
where $a$ is a positive constant.
Lemma 5. The support functions $\gamma$ of $M_{\gamma}$ in the above (I)-(IV) are respectively given by the followings:
(I) $\gamma\left(\nu_{3}\right)=\sqrt{1+b \nu_{3}^{2}}, \quad\left(b:=a^{2}-1>-1\right)$,
(II) $\gamma\left(\nu_{3}\right)=\sqrt{-1+b \nu_{3}^{2}}, \quad\left(b:=a^{2}+1>1, \frac{1}{\sqrt{b}}<\left|\nu_{3}\right| \leq\right.$ 1),
(III) $\gamma\left(\nu_{3}\right)=\sqrt{1-b \nu_{3}^{2}}, \quad\left(b:=a^{2}+1>1,\left|\nu_{3}\right|<\frac{1}{\sqrt{b}}\right)$,
(IV) $\gamma\left(\nu_{3}\right)=\frac{-1+\nu_{3}^{2}}{b \nu_{3}}, \quad\left(b:=4 a>0, \nu_{3} \neq 0\right)$.

Proof. (I) Represent the upper half of $M_{\gamma}$ as

$$
Y\left(x_{1}, x_{2}\right)=\left(x_{1}, x_{2}, a \sqrt{1-x_{1}^{2}-x_{2}^{2}}\right)
$$

The outward pointing unit normal $\nu$ to $Y$ is given by

$$
\nu=\frac{1}{\sqrt{1+\left(a^{2}-1\right)\left(x_{1}^{2}+x_{2}^{2}\right)}}\left(a x_{1}, a x_{2}, \sqrt{1-x_{1}^{2}-x_{2}^{2}}\right) .
$$

Hence, we obtain

$$
\gamma=\langle Y, \nu\rangle=\sqrt{1+\left(a^{2}-1\right) \nu_{3}^{2}}=\sqrt{1+b \nu_{3}^{2}}
$$

which proves (I).
Similarly, we obtain (II)-(IV).
Proposition 3. Let $\gamma$ be a function given by the above (I)-(IV). Then, there exist $\gamma$-minimal surfaces foliated by circles contained in parallel planes which are orthogonal to the $x_{3}$-axis. Up to translations in $\mathbf{R}^{3}$, rotations around the $x_{3}$-axis, and symmetry with respect to a plane $\left\{x_{3}=\right.$ constant $\}$, they are respectively represented as follows.
(I) Catenoid-type:

$$
\begin{equation*}
X(\theta, t)=\left(\frac{\cosh (c t)}{c \sqrt{1+b}} \cos \theta, \frac{\cosh (c t)}{c \sqrt{1+b}} \sin \theta, t\right), \quad c \neq 0 . \tag{22}
\end{equation*}
$$

Riemann-type:

$$
\begin{gather*}
X(\theta, r)=\left(r \cos \theta+\int \frac{c_{1} r^{2} d r}{\sqrt{c_{1}^{2} r^{4}+c_{2} r^{2}-1}}, r \sin \theta\right. \\
\left.\sqrt{1+b} \int \frac{d r}{\sqrt{c_{1}^{2} r^{4}+c_{2} r^{2}-1}}\right)  \tag{23}\\
c_{1} \neq 0, r \geq\left(\frac{-c_{2}+\sqrt{c_{2}^{2}+4 c_{1}^{2}}}{2 c_{1}^{2}}\right)^{1 / 2} \tag{24}
\end{gather*}
$$

(II) Catenoid-type:

$$
\begin{align*}
& X(\theta, t)=\left(\frac{\sinh (c t)}{c \sqrt{b-1}} \cos \theta, \frac{\sinh (c t)}{c \sqrt{b-1}} \sin \theta, t\right)  \tag{25}\\
& c \neq 0, t \neq 0 \tag{26}
\end{align*}
$$

## Riemann-type:

$$
\begin{align*}
& X(\theta, r)=\left(r \cos \theta+\int \frac{c_{1} r^{2} d r}{\sqrt{c_{1}^{2} r^{4}+c_{2} r^{2}+1}}, r \sin \theta\right. \\
& \left.\sqrt{b-1} \int \frac{d r}{\sqrt{c_{1}^{2} r^{4}+c_{2} r^{2}+1}}\right)  \tag{27}\\
& \quad c_{1} \neq 0, c_{2}>2\left|c_{1}\right|, r>0 \tag{28}
\end{align*}
$$

(III) Catenoid-type:

$$
\begin{array}{r}
X(\theta, t)=\left(\frac{\sin (c t)}{c \sqrt{b-1}} \cos \theta, \frac{\sin (c t)}{c \sqrt{b-1}} \sin \theta, t\right) \\
c \neq 0, \sin (c t) \neq 0 \tag{30}
\end{array}
$$

Riemann-type:

$$
\begin{gather*}
X(\theta, r)=\left(r \cos \theta+\int \frac{c_{1} r^{2} d r}{\sqrt{c_{1}^{2} r^{4}+c_{2} r^{2}+1}}, r \sin \theta,\right. \\
\left.\sqrt{b-1} \int \frac{d r}{\sqrt{c_{1}^{2} r^{4}+c_{2} r^{2}+1}}\right),  \tag{31}\\
c_{1} \neq 0, c_{2}<-2\left|c_{1}\right|, 0<r \leq\left(\frac{\left|c_{2}\right|-\sqrt{c_{2}^{2}-4 c_{1}^{2}}}{2 c_{1}^{2}}\right)^{1 / 2} . \tag{32}
\end{gather*}
$$

(IV) Catenoid-type:

$$
\begin{equation*}
X(\theta, t)=\left(e^{c t} \cos \theta, e^{c t} \sin \theta, t\right), \quad c \neq 0 \tag{33}
\end{equation*}
$$

## Riemann-type:

$$
\begin{align*}
X(\theta, r)= & \left(r \cos \theta+\int \frac{c_{1} r d r}{\sqrt{c_{1}^{2} r^{2}+c_{2}}}\right. \\
& \left.r \sin \theta, \int \frac{d r}{\sqrt{c_{1}^{2} r^{4}+c_{2} r^{2}}}\right) \tag{34}
\end{align*}
$$

$$
\begin{equation*}
c_{1} \neq 0, c_{2}>0, r>0 \tag{35}
\end{equation*}
$$



Figure 1: $\gamma$-minimal surfaces for $\gamma(\nu)$ as in (I) of Lemma 5.

(II) Catenoid-type

(II) Riemann-type

Figure 2: $\gamma$-minimal surfaces for $\gamma(\nu)$ as in (II) of Lemma 5.


Figure 3: $\gamma$-minimal surfaces for $\gamma(\nu)$ as in (III) of Lemma 5.

Remark 1. Let $\gamma$ be a function given in (I), (II), or (III) in Lemma 5. Set $\alpha:=\sqrt{b+1}$ for (I), and $\alpha:=\sqrt{b-1}$ for (II) and (III). If we take the transformation $\tilde{x}_{1}=x_{1}$, $\tilde{x}_{2}=x_{2}, \tilde{x}_{3}=x_{3} / \alpha$, then an immersion $X=\left(x_{1}, x_{2}, x_{3}\right)$ in $\mathbf{R}^{3}$ is $\gamma$-minimal if and only if $\tilde{X}=\left(\tilde{x}_{1}, \tilde{x}_{2}, \tilde{x}_{3}\right)$ is a minimal

(IV) Catenoid-type

(IV) Riemann-type

Figure 4: $\gamma$-minimal surfaces for $\gamma(\nu)$ as in (IV) of Lemma 5.
surface (for (I)), a spacelike zero mean curvature surface in $\mathbf{R}_{1}^{3}$ (for (II)), a timelike zero mean curvature surface in $\mathbf{R}_{1}^{3}$ (for (III)), respectively. This is proved by the same way as Example 4.4 in [8].
Remark 2. $\gamma$-minimal surfaces for $\gamma$ given in (IV) in Lemma 5 are graphs of harmonic functions (cf. [12]).
Proof of Proposition 3. Let $\mu_{1}, \mu_{2}$ be the principal curvatures of $M_{\gamma}$ with respect to the normal - $\nu$ in the same way as Lemma 2.

We represent the surface as

$$
\begin{equation*}
X(\theta, t)=(r(t) \cos \theta+f(t), r(t) \sin \theta+g(t), t) \tag{36}
\end{equation*}
$$

Note that $X$ is a surface of revolution around the $x_{3}$-axis if and only if $f \equiv g \equiv 0$ holds.

As in the proof of Lemma 4, we set

$$
\Delta:=r^{\prime}+f^{\prime} \cos \theta+g^{\prime} \sin \theta
$$

Then, the Gauss map $\nu$ of $X$ is

$$
\begin{aligned}
\nu(\theta, t) & :=\frac{X_{\theta} \times X_{t}}{\left|X_{\theta} \times X_{t}\right|} \\
& =\frac{1}{\left(\Delta^{2}+1\right)^{1 / 2}}(\cos \theta, \sin \theta,-\Delta)
\end{aligned}
$$

(I) By a simple computation using Lemma 2, we obtain

$$
\frac{1}{\mu_{1}}=\frac{b+1}{\left(1+b \nu_{3}^{2}\right)^{3 / 2}}, \quad \frac{1}{\mu_{2}}=\frac{1}{\sqrt{1+b \nu_{3}^{2}}}
$$

Since $\nu_{3}=\frac{-\Delta}{\left(\Delta^{2}+1\right)^{1 / 2}}$, we obtain

$$
\begin{align*}
& \frac{1}{\mu_{1}}=(b+1)\left(\frac{1+\Delta^{2}}{1+(b+1) \Delta^{2}}\right)^{3 / 2}  \tag{37}\\
& \frac{1}{\mu_{2}}=\left(\frac{1+\Delta^{2}}{1+(b+1) \Delta^{2}}\right)^{1 / 2} \tag{38}
\end{align*}
$$

By Lemma 4 with (37) and (38), we see that $\Lambda=0$ if and only if

$$
\begin{aligned}
& (b+1)\left(r f^{\prime \prime}-2 r^{\prime} f^{\prime}\right) \cos \theta+(b+1)\left(r g^{\prime \prime}-2 r^{\prime} g^{\prime}\right) \sin \theta \\
& \quad+(b+1)\left(r^{\prime \prime} r-\left(r^{\prime}\right)^{2}-\left(f^{\prime}\right)^{2}-\left(g^{\prime}\right)^{2}\right)-1=0
\end{aligned}
$$

holds. This gives the following system of ordinary differential equations:

$$
\begin{gather*}
r f^{\prime \prime}-2 r^{\prime} f^{\prime}=0  \tag{39}\\
r g^{\prime \prime}-2 r^{\prime} g^{\prime}=0  \tag{40}\\
(b+1)\left(r^{\prime \prime} r-\left(r^{\prime}\right)^{2}-\left(f^{\prime}\right)^{2}-\left(g^{\prime}\right)^{2}\right)-1=0 \tag{41}
\end{gather*}
$$

From (39) and (40), we have

$$
\begin{equation*}
f^{\prime}=c_{1} r^{2}, \quad g^{\prime}=c_{2} r^{2} \tag{42}
\end{equation*}
$$

First, assume $f^{\prime}=g^{\prime}=0$. Then, (41) is equivalent to

$$
\begin{equation*}
\left(r^{\prime \prime} r-\left(r^{\prime}\right)^{2}\right)-\frac{1}{b+1}=0 \tag{43}
\end{equation*}
$$

By a standard way, we see that the general solution of (43) is

$$
\begin{equation*}
r=\frac{\cosh \left(c_{3}\left(t+c_{4}\right)\right)}{c_{3} \sqrt{b+1}}, \quad c_{3} \neq 0 \tag{44}
\end{equation*}
$$

which gives the formula (22).
Next, we assume that $f^{\prime} \neq 0$ or $g^{\prime} \neq 0$ holds. Because of (42), $c_{2} f^{\prime}-c_{1} g^{\prime}=0$ holds. This implies that $c_{2} f-c_{1} g=$ constant. Since $\left(c_{1}, c_{2}\right) \neq(0,0)$, by rotating the surface around the $x_{3}$-axis if necessary, we may assume that

$$
\begin{equation*}
f^{\prime}=c_{1} r^{2}\left(c_{1} \neq 0\right), \quad g(t) \equiv 0 \tag{45}
\end{equation*}
$$

holds. Then, (41) is equivalent to

$$
1+(b+1)\left(\left(r^{\prime}\right)^{2}-r^{\prime \prime} r+c_{1}^{2} r^{4}\right)=0
$$

From this, by a standard argument, we obtain

$$
\begin{equation*}
\frac{d r}{d t}= \pm \sqrt{c_{1}^{2} r^{4}+c_{5} r^{2}-\frac{1}{b+1}} . \tag{46}
\end{equation*}
$$

Hence,

$$
\begin{align*}
t= & \pm \int \frac{d r}{\sqrt{c_{1}^{2} r^{4}+c_{5} r^{2}-\frac{1}{b+1}}} \\
= & \pm \sqrt{b+1} \int \frac{d r}{\sqrt{c_{6}^{2} r^{4}+c_{7} r^{2}-1}}  \tag{47}\\
& c_{6}:=\sqrt{b+1} c_{1} \neq 0, \quad c_{7}:=(b+1) c_{5}
\end{align*}
$$

By using (45) and (46), we easily obtain

$$
\begin{equation*}
f= \pm \int \frac{c_{6} r^{2} d r}{\sqrt{c_{6}^{2} r^{4}+c_{7} r^{2}-1}}, \quad c_{6} \neq 0 \tag{48}
\end{equation*}
$$

(47) with (48) gives the formula (23). Moreover,

$$
c_{1}^{2} r^{4}+c_{2} r^{2}-1 \geq 0
$$

if and only if

$$
r \geq\left(\frac{-c_{2}+\sqrt{c_{2}^{2}+4 c_{1}^{2}}}{2 c_{1}^{2}}\right)^{1 / 2}
$$

holds, which gives the condition (24).
(II) The proof is similar to the proof of (I). We have

$$
\begin{aligned}
\frac{1}{\mu_{1}} & =\frac{-(b-1)}{\left(-1+b \nu_{3}^{2}\right)^{3 / 2}} \\
& =-(b-1)\left(\frac{1+\Delta^{2}}{-1+(b-1) \Delta^{2}}\right)^{3 / 2}
\end{aligned}
$$

$$
\frac{1}{\mu_{2}}=\frac{-1}{\sqrt{-1+b \nu_{3}^{2}}}=-\left(\frac{1+\Delta^{2}}{-1+(b-1) \Delta^{2}}\right)^{1 / 2}
$$

We see that $\Lambda=0$ if and only if

$$
\begin{aligned}
& (b-1)\left(r f^{\prime \prime}-2 r^{\prime} f^{\prime}\right) \cos \theta+(b-1)\left(r g^{\prime \prime}-2 r^{\prime} g^{\prime}\right) \sin \theta \\
& \quad+(b-1)\left\{r^{\prime \prime} r-\left(r^{\prime}\right)^{2}-\left(f^{\prime}\right)^{2}-\left(g^{\prime}\right)^{2}\right\}+1=0
\end{aligned}
$$

holds. This gives the following system of ordinary differential equations:

$$
\begin{gather*}
r f^{\prime \prime}-2 r^{\prime} f^{\prime}=0  \tag{49}\\
r g^{\prime \prime}-2 r^{\prime} g^{\prime}=0  \tag{50}\\
(b-1)\left(r^{\prime \prime} r-\left(r^{\prime}\right)^{2}-\left(f^{\prime}\right)^{2}-\left(g^{\prime}\right)^{2}\right)+1=0 \tag{51}
\end{gather*}
$$

First, assume $f^{\prime}=g^{\prime}=0$. Then, (51) is equivalent to

$$
\begin{equation*}
\left(r^{\prime \prime} r-\left(r^{\prime}\right)^{2}\right)+\frac{1}{b-1}=0 \tag{52}
\end{equation*}
$$

We have the following three types of general solutions of (52):

$$
\begin{align*}
& r=\frac{1}{\sqrt{b-1}} t+c  \tag{53}\\
& r=\frac{\sinh \left(c_{1}\left(t+c_{2}\right)\right)}{c_{1} \sqrt{b-1}}, \quad c_{1} \neq 0  \tag{54}\\
& r=\frac{\sin \left(c_{1}\left(t+c_{2}\right)\right)}{c_{1} \sqrt{b-1}}, \quad c_{1} \neq 0 \tag{55}
\end{align*}
$$

By a suitable translation, the corresponding surfaces are given by

$$
\begin{align*}
& X(\theta, t)=\left(\frac{t}{\sqrt{b-1}} \cos \theta, \frac{t}{\sqrt{b-1}} \sin \theta, t\right)  \tag{56}\\
& X(\theta, t)=\left(\frac{\sinh (c t)}{c \sqrt{b-1}} \cos \theta, \frac{\sinh (c t)}{c \sqrt{b-1}} \sin \theta, t\right), c \neq 0  \tag{57}\\
& X(\theta, t)=\left(\frac{\sin (c t)}{c \sqrt{b-1}} \cos \theta, \frac{\sin (c t)}{c \sqrt{b-1}} \sin \theta, t\right), c \neq 0 \tag{58}
\end{align*}
$$

respectively. Later, we will show that in order that the surface is compatible with $\gamma$, the surface must be given by (57), which gives the formula (25).

Next, we assume that $f^{\prime} \neq 0$ or $g^{\prime} \neq 0$ holds. By rotating the surface around the $x_{3}$-axis if necessary, we may assume that

$$
\begin{equation*}
f^{\prime}=c_{1} r^{2}\left(c_{1} \neq 0\right), \quad g(t) \equiv 0 \tag{59}
\end{equation*}
$$

holds. Then, (51) is equivalent to

$$
\begin{equation*}
1-(b-1)\left(\left(r^{\prime}\right)^{2}-r^{\prime \prime} r+c_{1}^{2} r^{4}\right)=0 . \tag{60}
\end{equation*}
$$

From this, by a standard argument, we obtain

$$
\begin{equation*}
\frac{d r}{d t}= \pm \sqrt{c_{1}^{2} r^{4}+c_{2} r^{2}+\frac{1}{b-1}} . \tag{61}
\end{equation*}
$$

Hence,

$$
\begin{align*}
& t= \pm \sqrt{b-1} \int \frac{d r}{\sqrt{c_{3}^{2} r^{4}+c_{4} r^{2}+1}}  \tag{62}\\
& \quad c_{3}:=\sqrt{b-1} c_{1} \neq 0, c_{4}:=(b-1) c_{2} .
\end{align*}
$$

By using $f^{\prime}=c_{1} r^{2}$ and (61), we obtain

$$
\begin{equation*}
f= \pm \int \frac{c_{3} r^{2} d r}{\sqrt{c_{3}^{2} r^{4}+c_{4} r^{2}+1}}, \quad c_{3} \neq 0 \tag{63}
\end{equation*}
$$

(62) with (63) gives the formula (27).

Next, we will check whether the surface is compatible with $\gamma$ or not. Note that the surface $X$ is compatible with $\gamma$ if and only if its Gauss map $\nu=\left(\nu_{1}, \nu_{2}, \nu_{3}\right)$ satisfies $b \nu_{3}^{2}-1>0$ for all $\theta$.

For the surface $X$ given by (56), the Gauss map $\nu$ is

$$
\begin{aligned}
\nu & :=\frac{X_{\theta} \times X_{t}}{\left|X_{\theta} \times X_{t}\right|} \\
& =b^{-1 / 2}(\sqrt{b-1} \cos \theta, \sqrt{b-1} \sin \theta,-1) .
\end{aligned}
$$

This shows that $b \nu_{3}^{2}-1 \equiv 0$, and hence $X$ is not compatible with $\gamma$.

For the surface $X$ given by (57),

$$
\nu=\left(1+\frac{\cosh ^{2}(c t)}{b-1}\right)^{-1 / 2}\left(\cos \theta, \sin \theta,-\frac{\cosh (c t)}{\sqrt{b-1}}\right) .
$$

This shows that $b \nu_{3}^{2}-1 \geq 0$ always holds, and that $b \nu_{3}^{2}-1>$ 0 for all $\theta$ if and only if $t \neq 0$.

For the surface $X$ given by (58),

$$
\begin{equation*}
\nu=\left(1+\frac{\cos ^{2}(c t)}{b-1}\right)^{-1 / 2}\left(\cos \theta, \sin \theta,-\frac{\cos (c t)}{\sqrt{b-1}}\right) . \tag{64}
\end{equation*}
$$

This shows that $b \nu_{3}^{2}-1 \leq 0$ always holds, and hence $X$ is not compatible with $\gamma$.

For the surface $X$ given by (27), the Gauss map $\nu$ is

$$
\begin{align*}
\nu:= & \frac{X_{\theta} \times X_{r}}{\left|X_{\theta} \times X_{r}\right|}  \tag{65}\\
X_{\theta} \times X_{r}= & \left(\frac{\sqrt{b-1} r \cos \theta}{\left(c_{1}^{2} r^{4}+c_{2} r^{2}+1\right)^{1 / 2}}, \frac{\sqrt{b-1} r \sin \theta}{\left(c_{1}^{2} r^{4}+c_{2} r^{2}+1\right)^{1 / 2}}\right. \\
& \left.-r-\frac{c_{1} r^{3} \cos \theta}{\left(c_{1}^{2} r^{4}+c_{2} r^{2}+1\right)^{1 / 2}}\right) . \tag{66}
\end{align*}
$$

This shows that, $b \nu_{3}^{2}-1>0$ for all $\theta$ if and only if $c_{2}>2\left|c_{1}\right|$ holds.
(III) The proof is again similar to the proof of (I). We see that the condition $\Lambda=0$ is equivalent to the condition that the system of ordinary differential equations (49), (50) and (51) holds.
First, assume $f^{\prime}=g^{\prime}=0$. Then, (51) is equivalent to (52). Note that $b-1>0$. The general solutions of (52) are given by (53), (54), and (55), and corresponding surfaces are given by (56), (57), and (58). Note that the surface $X$ is compatible with $\gamma$ if and only if its Gauss map $\nu=\left(\nu_{1}, \nu_{2}, \nu_{3}\right)$ satisfies $1-b \nu_{3}^{2}>0$ for all $\theta$. As we have seen above, for the surfaces (56) and (57), $b \nu_{3}^{2}-1 \geq 0$ holds at every point. Hence, they are not compatible with $\gamma$. Therefore, only the possibility is the case (58), which is the same as the formula (29). The Gauss map $\nu$ for this
surface is given by (64), which shows that, $1-b \nu_{3}^{2}>0$ for all $\theta$ if and only if $\sin (c t) \neq 0$.
Next, we assume that $f^{\prime} \neq 0$ or $g^{\prime} \neq 0$ holds. By rotating the surface around the $x_{3}$-axis if necessary, we may assume that $f$ and $g$ satisfy (59). Then, (51) is equivalent to (60). Hence, we obtain the formula (31).

We will check whether the surface is compatible with $\gamma$ or not. The Gauss map $\nu$ is given by (65), (66) as in the case (II). This shows that, $1-b \nu_{3}^{2}>0$ for all $\theta$ if and only if both $c_{2}<-2\left|c_{1}\right|$ and

$$
0<r \leq\left(\frac{\left|c_{2}\right|-\sqrt{c_{2}^{2}-4 c_{1}^{2}}}{2 c_{1}^{2}}\right)^{1 / 2}
$$

hold, which gives the condition (32).
(IV) The proof is again similar to the proof of (I). We obtain

$$
\begin{align*}
& \frac{1}{\mu_{1}}=\frac{-2}{b \nu_{3}^{3}}=\frac{2\left(\Delta^{2}+1\right)^{3 / 2}}{b \Delta^{3}}  \tag{67}\\
& \frac{1}{\mu_{2}}=\frac{-2}{b \nu_{3}}=\frac{2\left(\Delta^{2}+1\right)^{1 / 2}}{b \Delta} \tag{68}
\end{align*}
$$

By Lemma 4 with (67) and (68), we see that $\Lambda=0$ if and only if

$$
\begin{aligned}
\left(r f^{\prime \prime}-2 r^{\prime} f^{\prime}\right) \cos \theta & +\left(r g^{\prime \prime}-2 r^{\prime} g^{\prime}\right) \sin \theta \\
& +\left(r^{\prime \prime} r-\left(r^{\prime}\right)^{2}-\left(f^{\prime}\right)^{2}-\left(g^{\prime}\right)^{2}\right)=0
\end{aligned}
$$

holds. This gives the following system of ordinary differential equations:

$$
\begin{gather*}
r f^{\prime \prime}-2 r^{\prime} f^{\prime}=0  \tag{69}\\
r g^{\prime \prime}-2 r^{\prime} g^{\prime}=0  \tag{70}\\
r^{\prime \prime} r-\left(r^{\prime}\right)^{2}-\left(f^{\prime}\right)^{2}-\left(g^{\prime}\right)^{2}=0 \tag{71}
\end{gather*}
$$

From (69) and (70), we have

$$
\begin{equation*}
f^{\prime}=c_{1} r^{2}, \quad g^{\prime}=c_{2} r^{2} . \tag{72}
\end{equation*}
$$

When $f^{\prime}=g^{\prime}=0$ holds, (71) is equivalent to

$$
\begin{equation*}
r^{\prime \prime} r-\left(r^{\prime}\right)^{2}=0 \tag{73}
\end{equation*}
$$

The general solution of (73) is

$$
r=e^{c_{1} t+c_{2}}
$$

which gives the formula (33).
When $f^{\prime} \neq 0$ or $g^{\prime} \neq 0$ holds, by rotating the surface around the $x_{3}$-axis if necessary, we may assume that

$$
\begin{equation*}
f^{\prime}=c_{1} r^{2}\left(c_{1} \neq 0\right), \quad g(t) \equiv 0 \tag{74}
\end{equation*}
$$

holds. Then, (71) is equivalent to

$$
\begin{equation*}
\left(r^{\prime}\right)^{2}-r^{\prime \prime} r+c_{1}^{2} r^{4}=0 \tag{75}
\end{equation*}
$$

From this, by a standard argument, we obtain

$$
\begin{equation*}
\frac{d r}{d t}= \pm \sqrt{c_{1}^{2} r^{4}+c_{2} r^{2}} \tag{76}
\end{equation*}
$$

Hence,

$$
\begin{equation*}
t= \pm \int \frac{d r}{\sqrt{c_{1}^{2} r^{4}+c_{2} r^{2}}} \tag{77}
\end{equation*}
$$

By using $f^{\prime}=c_{1} r^{2}$ and (76), we obtain

$$
\begin{equation*}
f= \pm \int \frac{c_{1} r d r}{\sqrt{c_{1}^{2} r^{2}+c_{2}}}, \quad c_{1} \neq 0 \tag{78}
\end{equation*}
$$

(77) with (78) gives the formula (34).

Next, we will check whether the surface is compatible with $\gamma$ or not. Note that the surface $X$ is compatible with $\gamma$ if and only if its Gauss map $\nu=\left(\nu_{1}, \nu_{2}, \nu_{3}\right)$ satisfies $\nu_{3} \neq 0$ for all $\theta$.

For the surface $X$ given by (33), the Gauss map $\nu$ is

$$
\begin{aligned}
\nu & :=\frac{X_{\theta} \times X_{t}}{\left|X_{\theta} \times X_{t}\right|} \\
& =\left(1+c^{2} e^{2 c t}\right)^{-1 / 2}\left(\cos \theta, \sin \theta,-c e^{c t}\right)
\end{aligned}
$$

This shows that, $\nu_{3} \neq 0$ for all $\theta$ if and only if $c \neq 0$.
For the surface $X$ given by (34), the Gauss map $\nu$ is

$$
\begin{aligned}
\nu:= & \frac{X_{\theta} \times X_{r}}{\left|X_{\theta} \times X_{r}\right|} \\
X_{\theta} \times X_{r}= & \left(\frac{\cos \theta}{\left(c_{1}^{2} r^{2}+c_{2}\right)^{1 / 2}}, \frac{\sin \theta}{\left(c_{1}^{2} r^{2}+c_{2}\right)^{1 / 2}}\right. \\
& \left.-r\left(1+\frac{c_{1} r \cos \theta}{\left(c_{1}^{2} r^{2}+c_{2}\right)^{1 / 2}}\right)\right)
\end{aligned}
$$

This shows that, $\nu_{3} \neq 0$ for all $\theta$ if and only if $c_{2}>0$ holds.

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